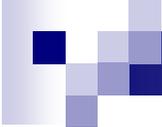




VUV & SX Optics

Kenta Amemiya

Photon Factory, High Energy Accelerator Research Organization



Outline

1. Mirrors

1.1. focusing & collimation

1.2. substrate materials

1.3. coating & reflectivity

2. Diffraction Gratings

2.1. principle - wavelength dispersion -

2.2. energy resolution

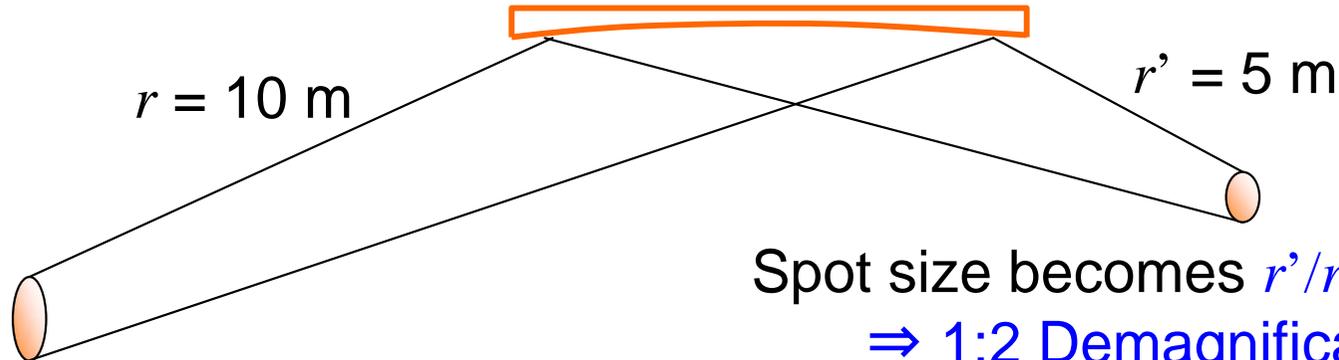
- dispersion and focus -

2.3. fabrication of gratings

1.1. focusing & collimation

Demagnification & Aberration

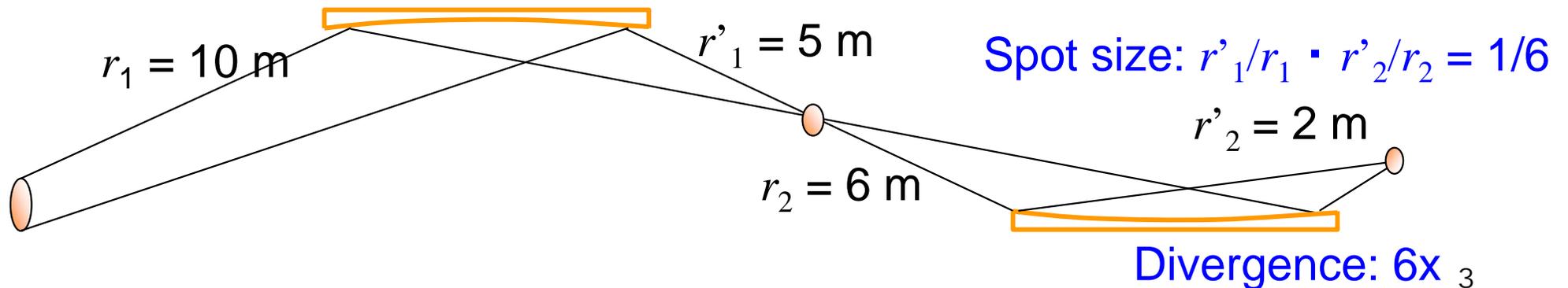
$$R = \frac{2}{(1/r + 1/r') \cos \theta}$$



Spot size becomes $r'/r = 1/2$
 \Rightarrow 1:2 Demagnification
 * Divergence becomes 2 times larger

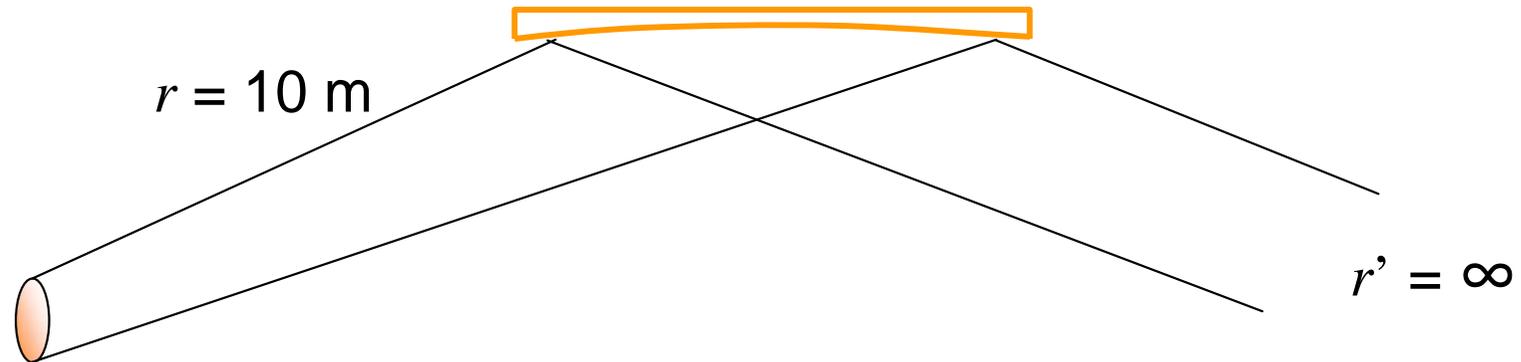
Elliptical mirror \Rightarrow perfect focus (determined by demagnification only)
 Use of a cylindrical or spherical mirror \Rightarrow poor focus (aberration)
 Larger divergence (larger illumination area) \Rightarrow larger aberration

Combination of focusing mirrors



1.1. focusing & collimation

Collimation

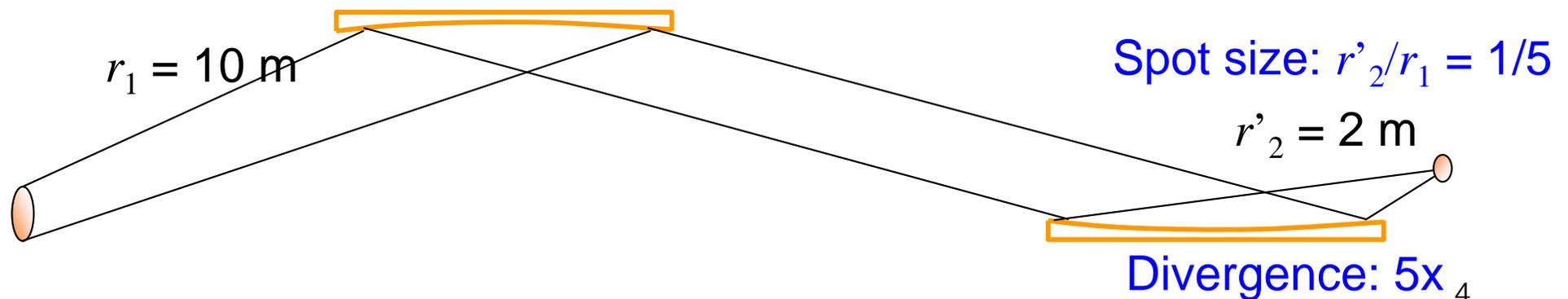


Parabolic mirror \Rightarrow perfect collimation (spot-size limited)

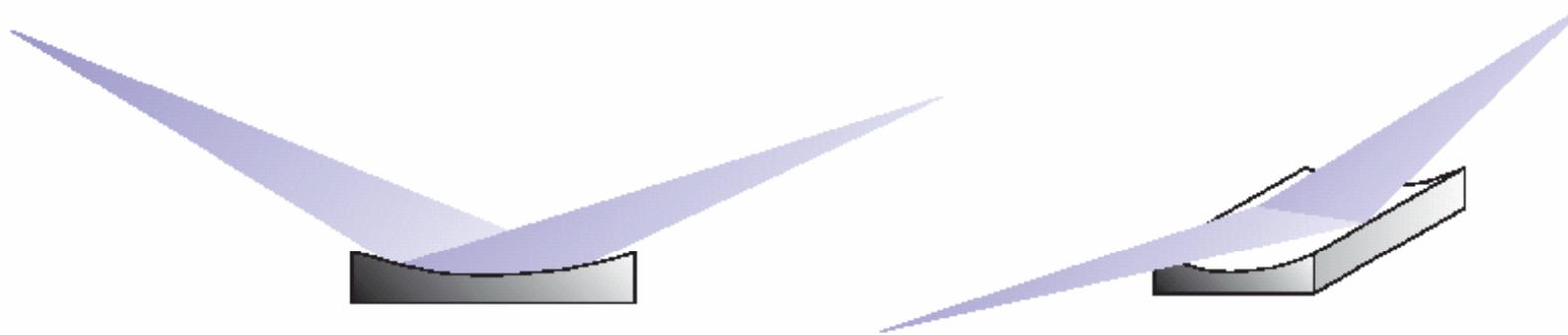
Use of a cylindrical or spherical mirror \Rightarrow poor collimation (aberration)

Larger divergence (larger illumination area) \Rightarrow larger aberration

Combination of collimating mirrors



1.1. focusing & collimation



Tangential focusing

Sagittal focusing

Aberration: smaller in **tangential** focusing

* Sagittal focusing is often adopted undulator beamlines

Radius: smaller in **sagittal** focusing

e.g. 88 deg (from normal),

$r = 10$ m, $r' = 5$ m

⇒ $R = 190$ m (tangential focusing), $\rho = 0.23$ m (sagittal focusing)

$$R = \frac{2}{(1/r + 1/r') \cos \theta}$$

$$\rho = \frac{2 \cos \theta}{(1/r + 1/r')}$$

Higher precision in **sagittal** focusing

e.g. $R = 190$ m $\pm 3\%$, $\rho = 0.23$ m $\pm < 0.5\%$

Slope error effects: smaller in **sagittal** focusing

1.2. substrate materials

Needs:

Easy to fabricate: precise control of the mirror shape

Low defects, pores

Hardness: small distortion

High thermal conductivity: cooling efficiency

Low thermal expansion: against a heat load

Typical materials

Si: for high heat load, with cooling

SiO₂: without cooling

* SiO₂ is suitable for mirror current measurements

1.3. coating & reflectivity

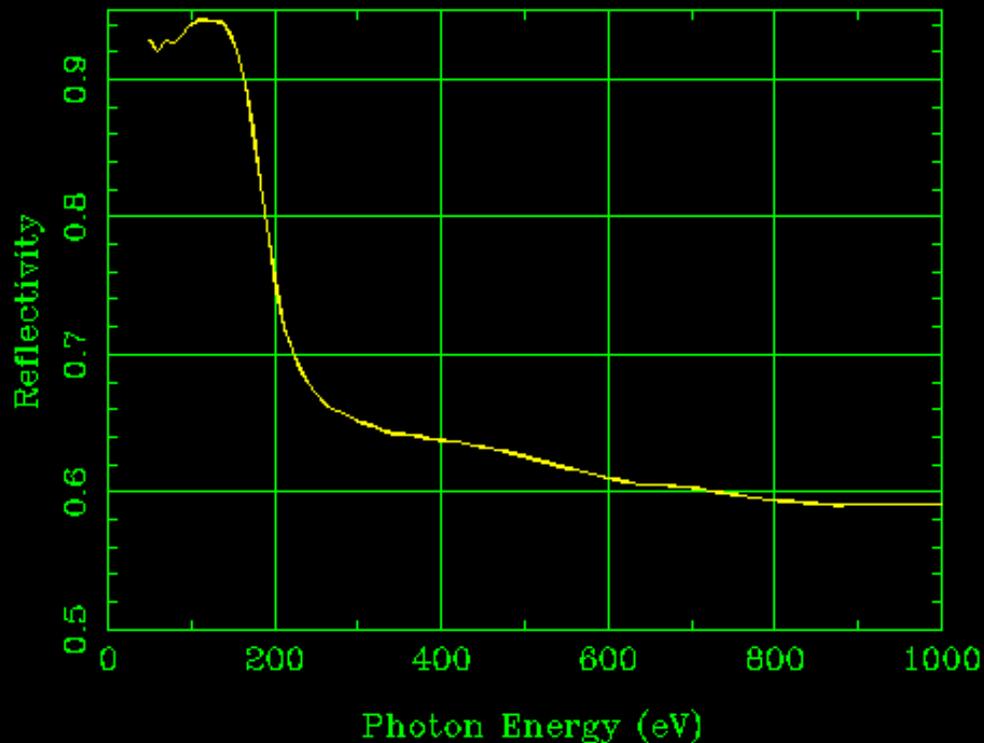
Coating layer:

High reflectivity by “total reflection”

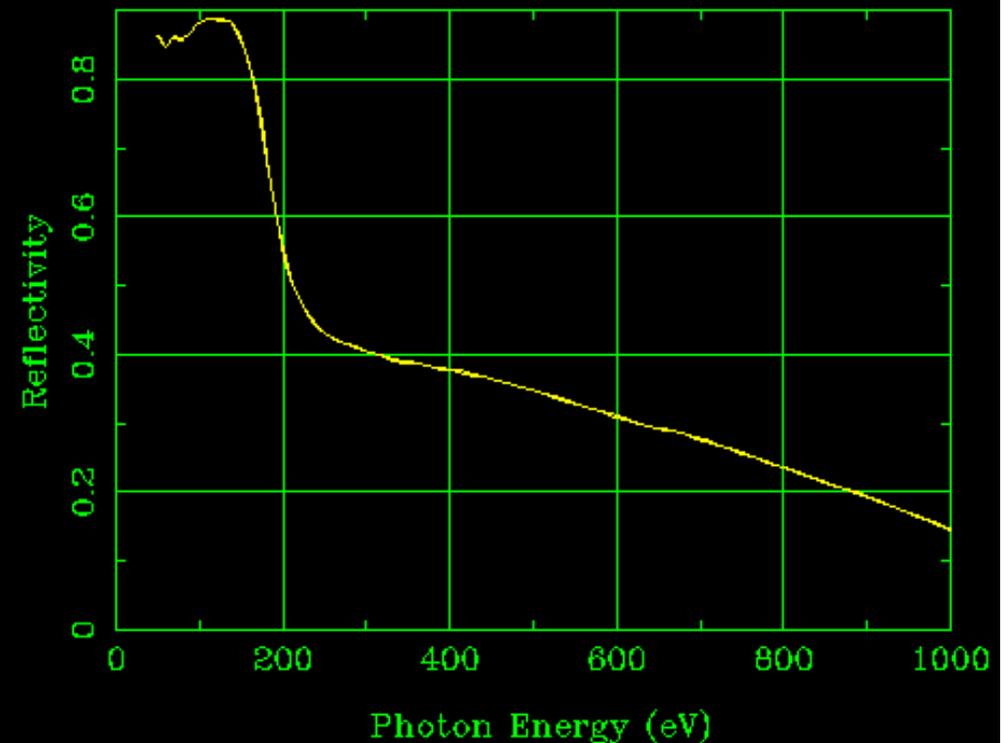
Protection against oxidation etc.

Typical coating material in VUV-SX: Au (50-100 nm)

88 deg



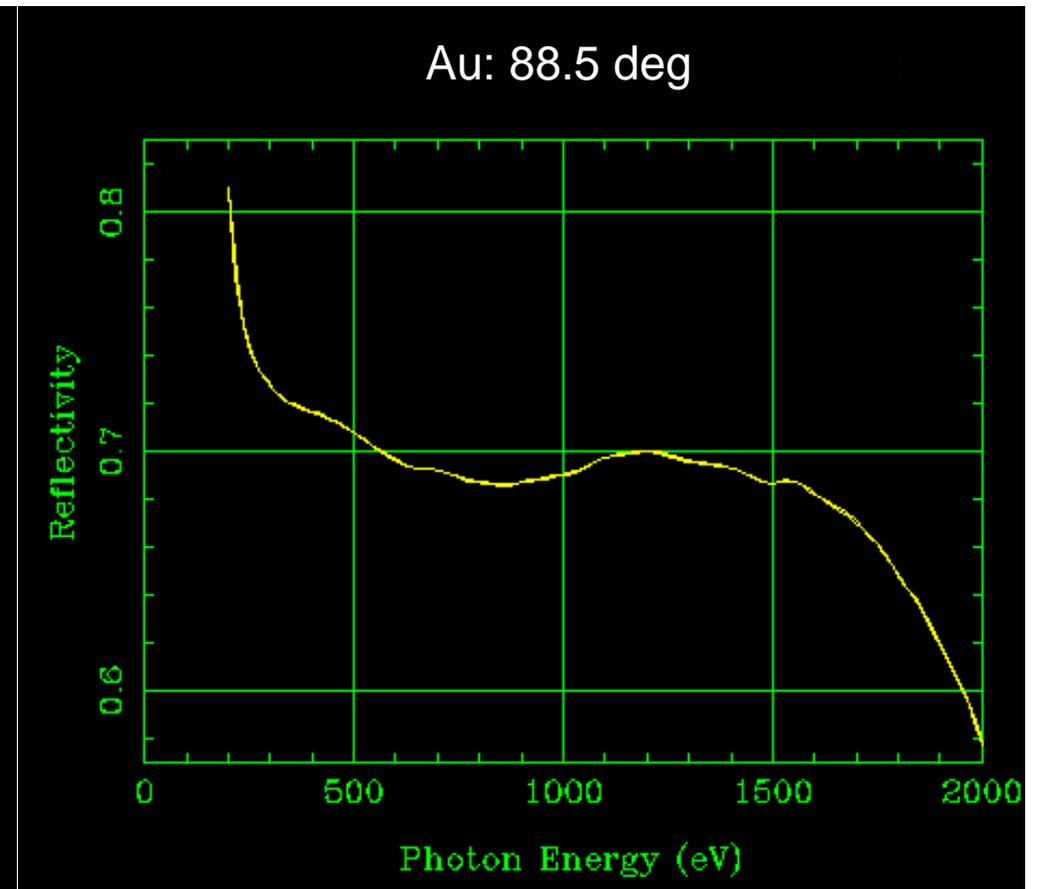
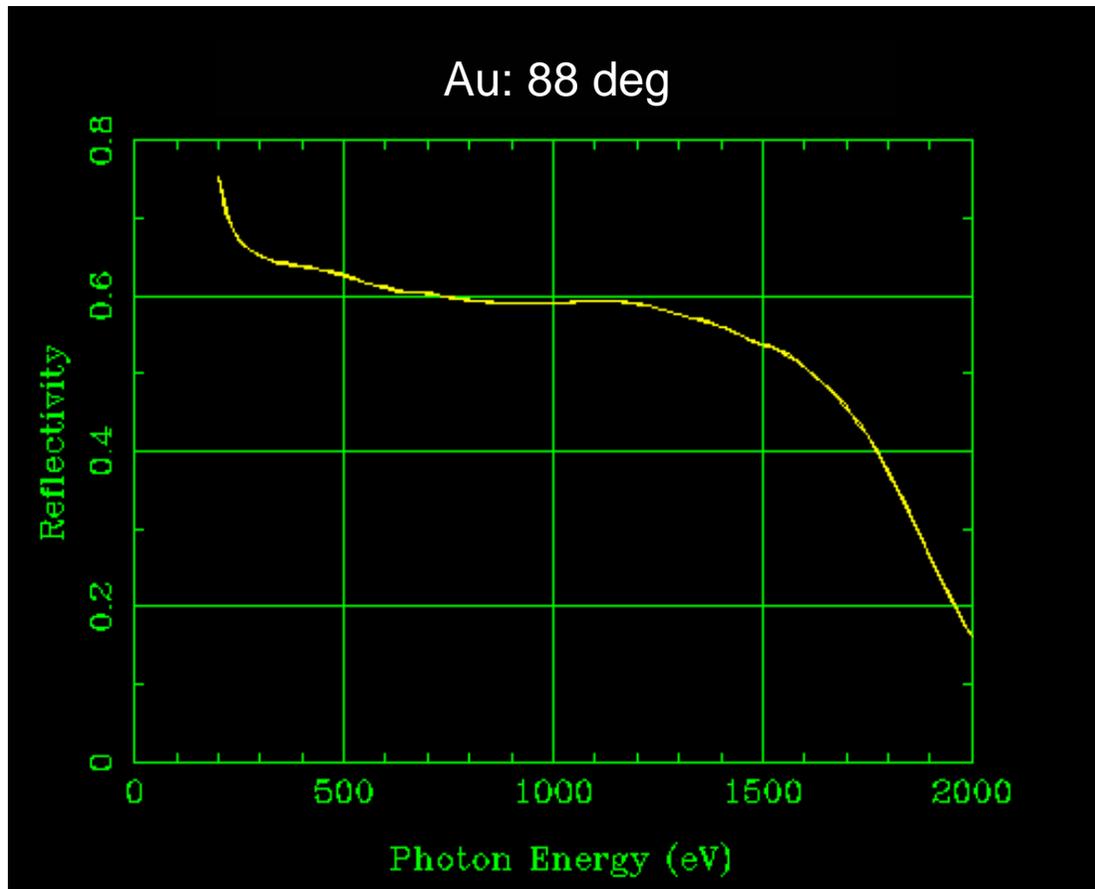
86 deg



1.3. coating & reflectivity

Higher energy region

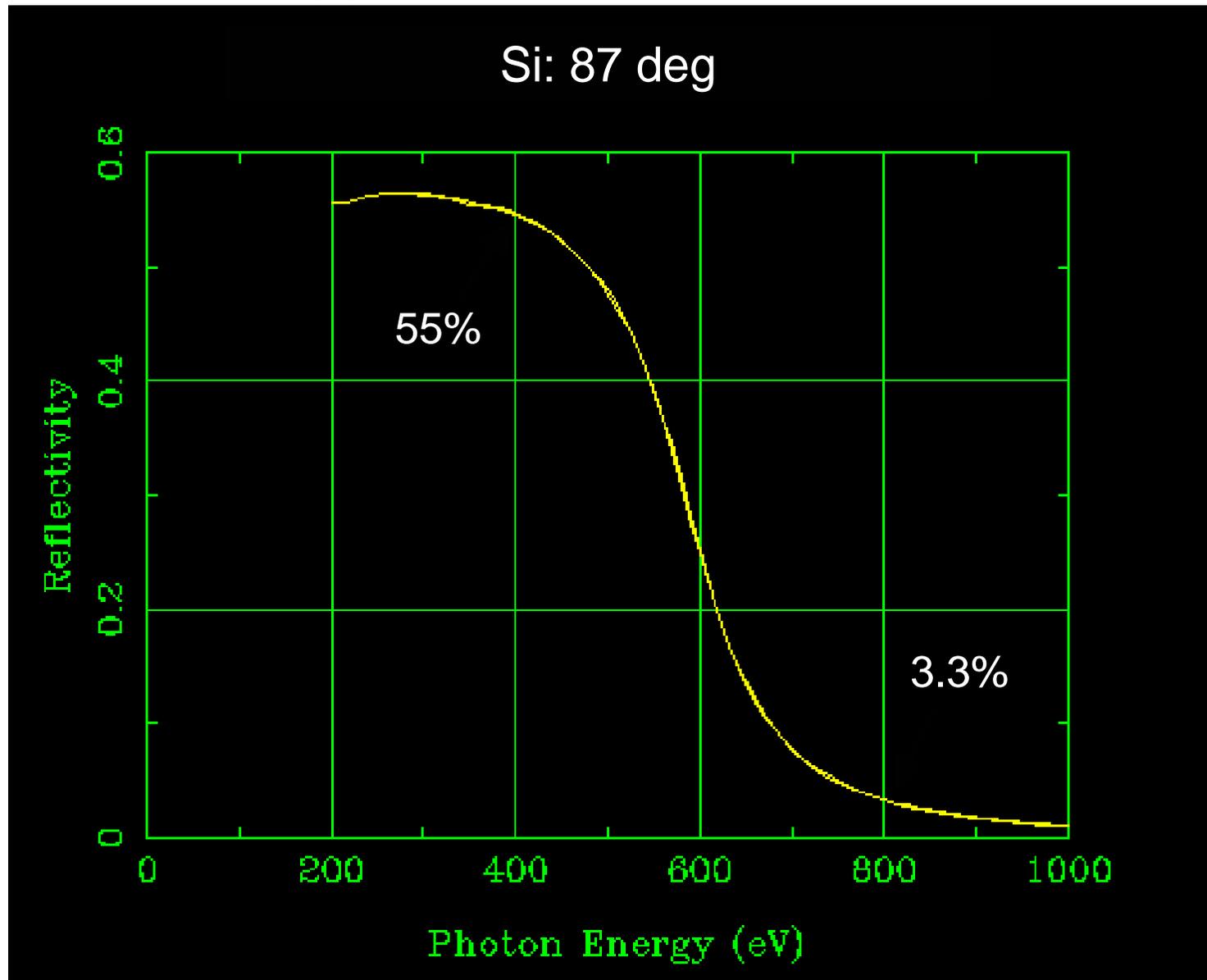
Critical energy for total reflection depends on incidence angle



http://henke.lbl.gov/optical_constants/

1.3. coating & reflectivity

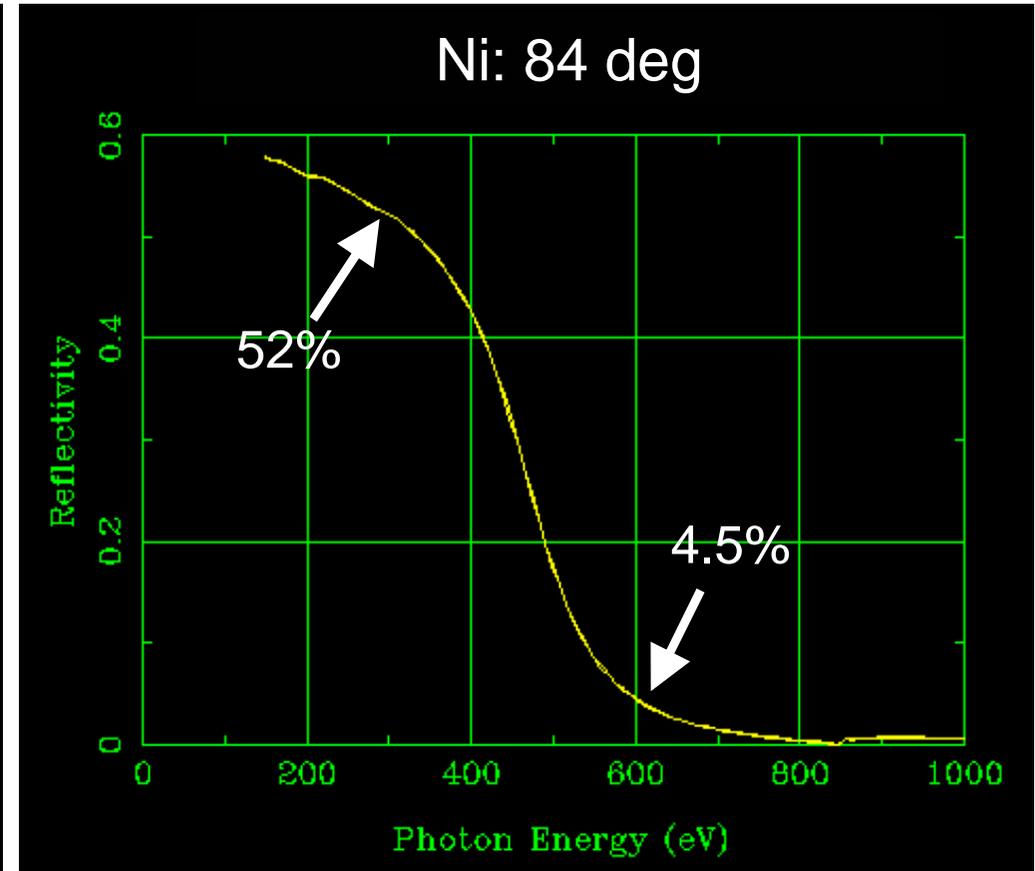
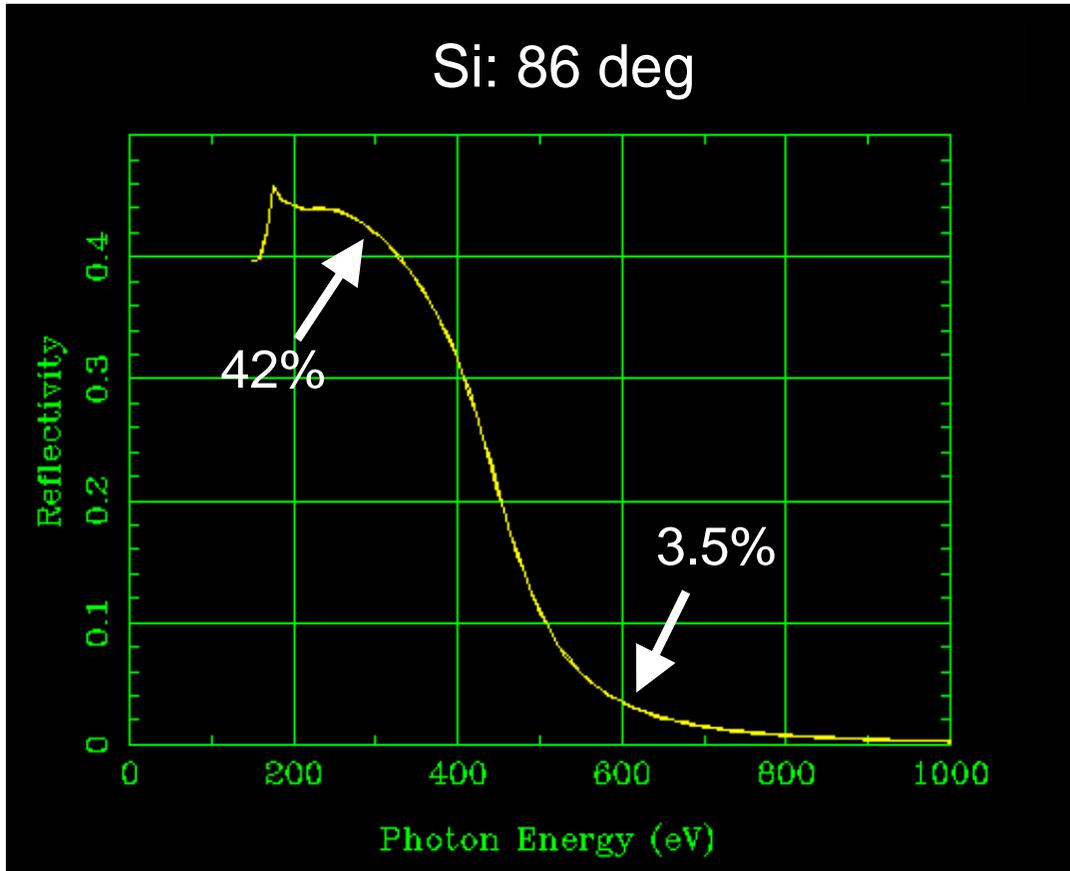
Higher order suppression



http://henke.lbl.gov/optical_constants/

1.3. coating & reflectivity

Higher order suppression: coating material dependence

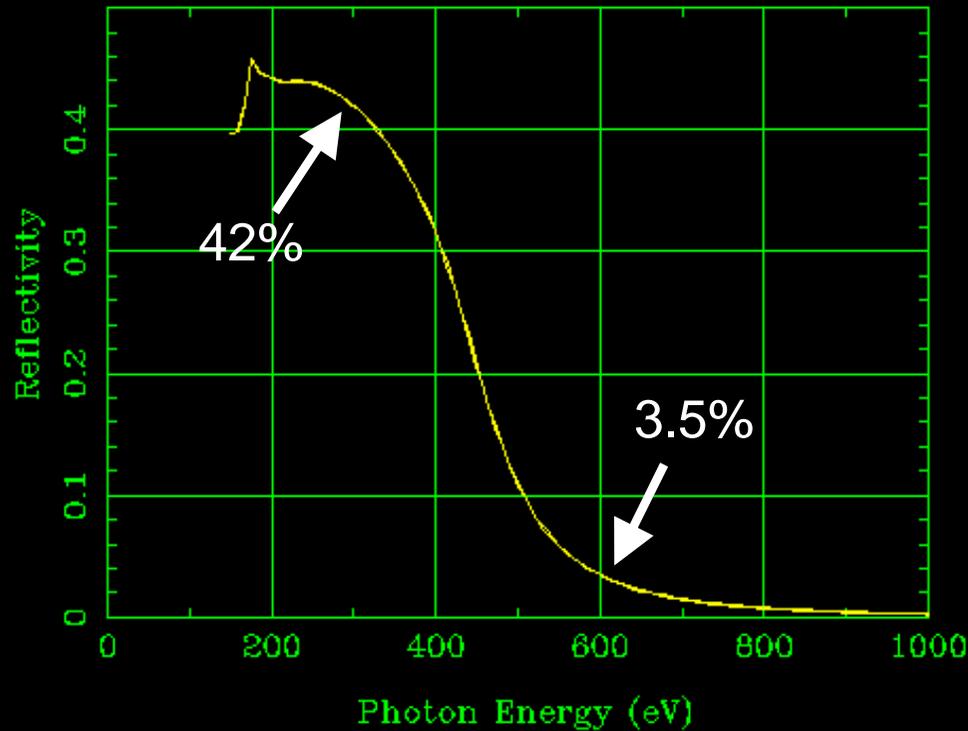


http://henke.lbl.gov/optical_constants/

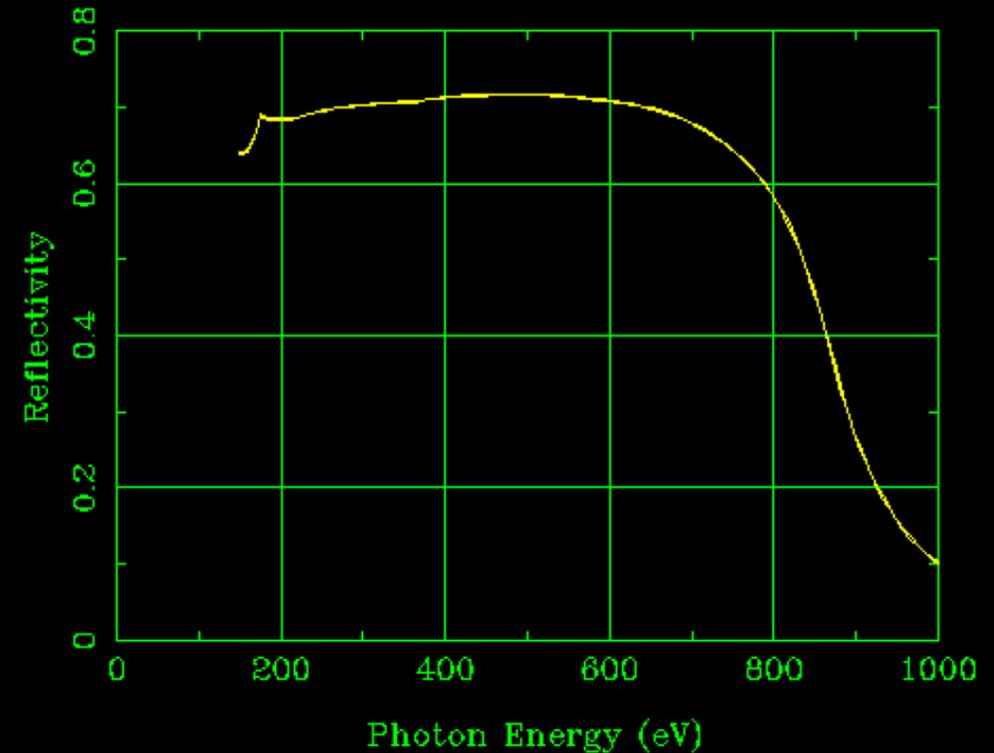
1.3. coating & reflectivity

Higher order suppression: **incidence-angle dependence**

Si: 86 deg

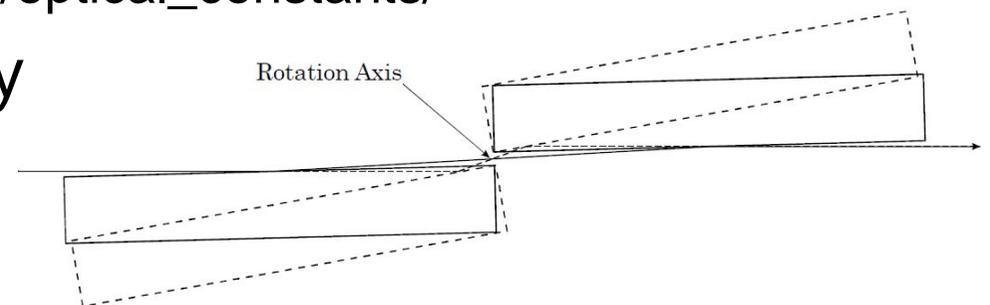


Si: 88 deg



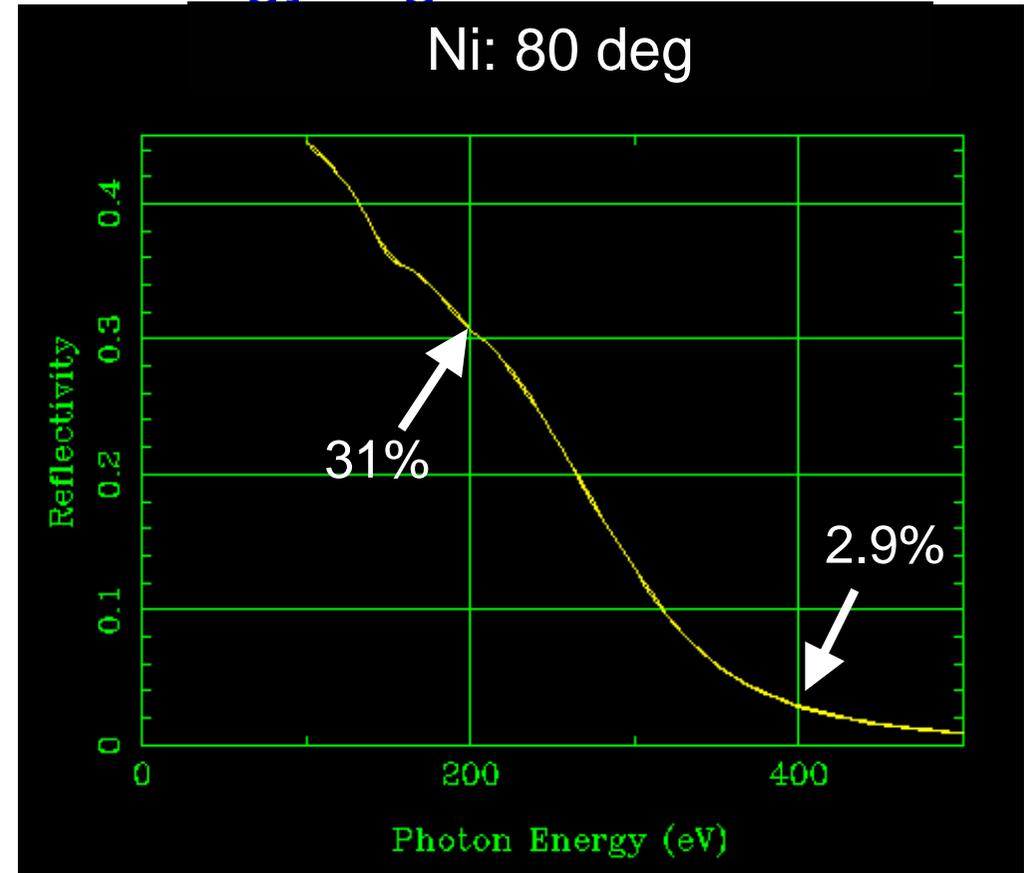
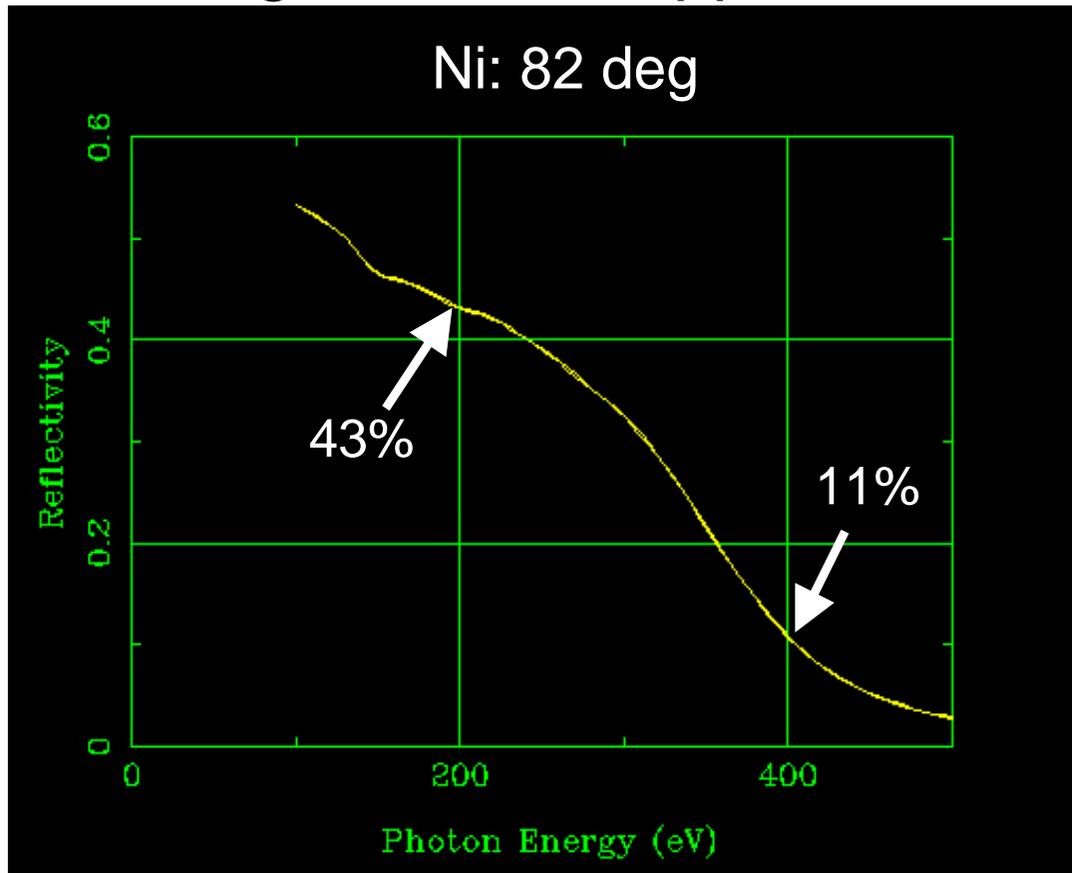
http://henke.lbl.gov/optical_constants/

Simultaneous rotation with energy scan is necessary
No precise control is required



1.3. coating & reflectivity

Higher order suppression in **low energy region**



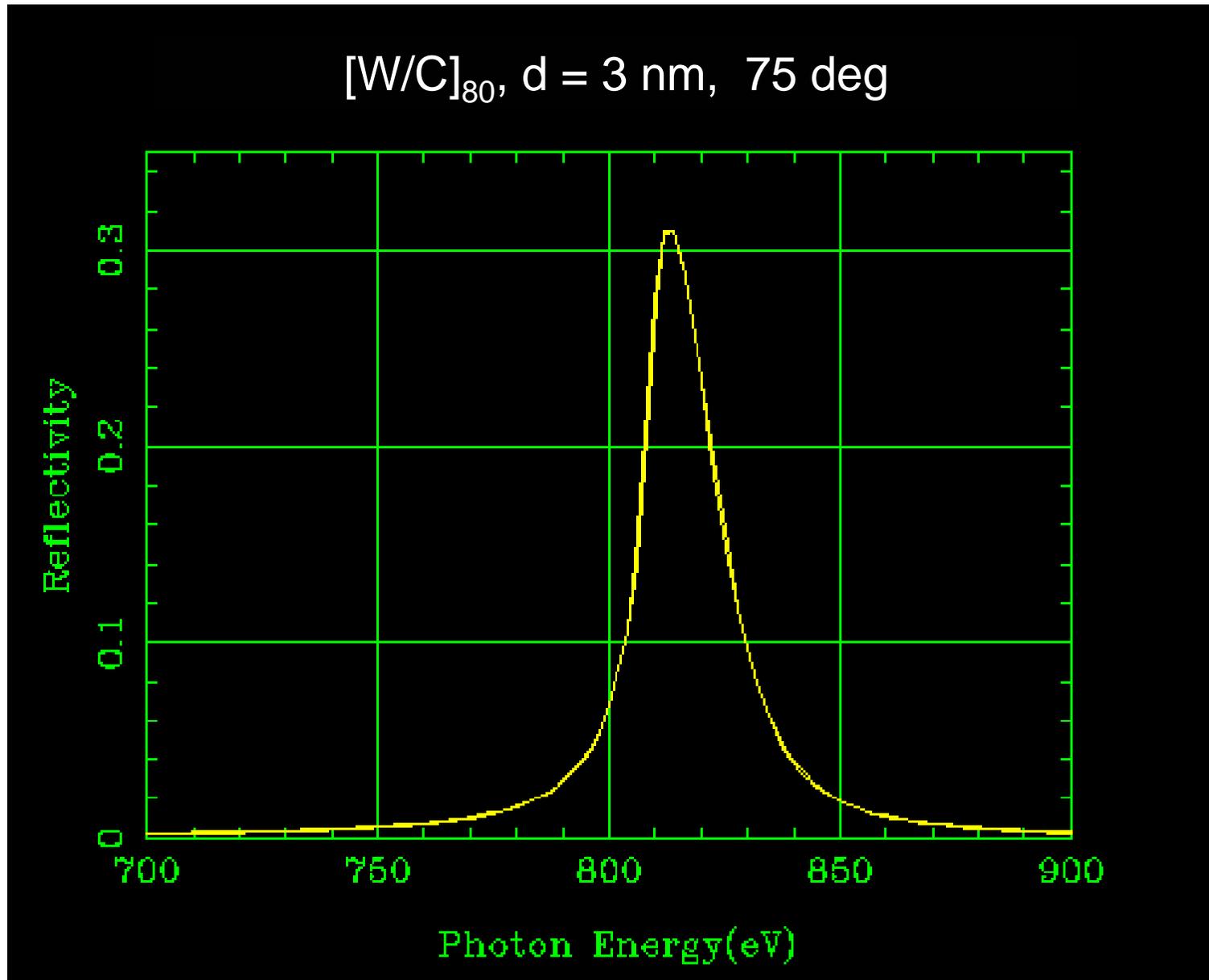
http://henke.lbl.gov/optical_constants/

It is difficult to achieve high reduction ratio
keeping high reflectivity for fundamental light

Not effective below ~100 eV

1.3. coating & reflectivity

Multilayer mirror





1. Mirrors

1.1. focusing & collimation

1.2. substrate materials

1.3. coating & reflectivity

2. Diffraction Gratings

2.1. principle - wavelength dispersion -

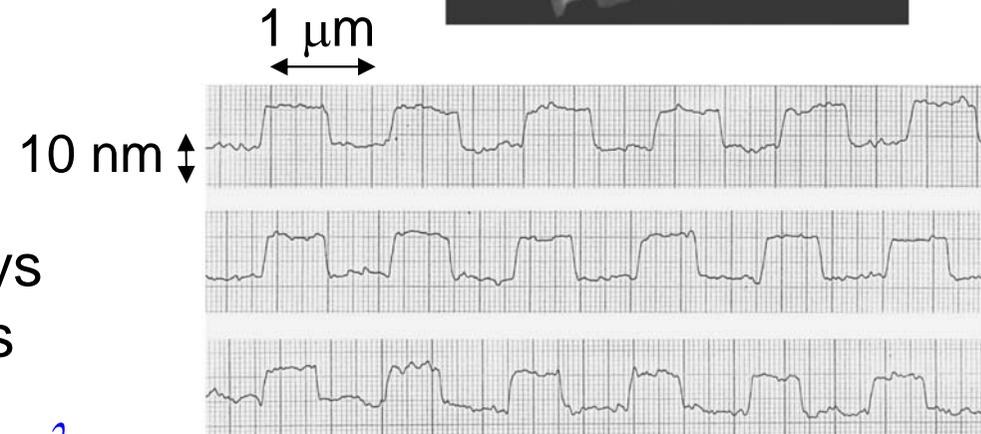
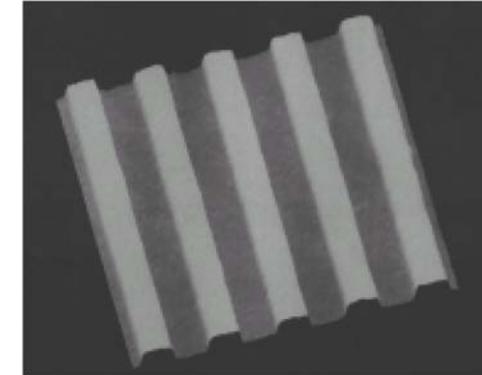
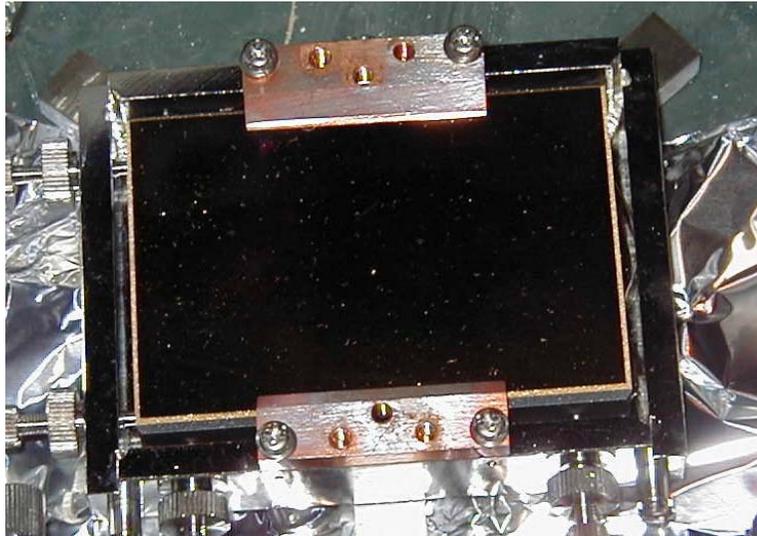
2.2. energy resolution

- dispersion and focus -

2.3. fabrication of gratings

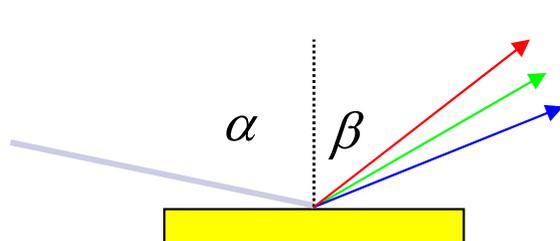
2.1. principle - wavelength dispersion -

Diffraction grating: **Periodic grooves** on a substrate



Principle: Interference between the rays reflected at different grooves

Enhanced when **light path difference** = $m\lambda$



* $\alpha > 0$

$$\sin \alpha + \sin \beta = nm\lambda$$

β depends on $\lambda \Rightarrow$ **Wavelength dispersion**

* $\alpha \neq |\beta|$

(if $\alpha = |\beta|$, any λ satisfies the above condition at $m = 0$)

\Rightarrow **zero-th order light**

n : groove density

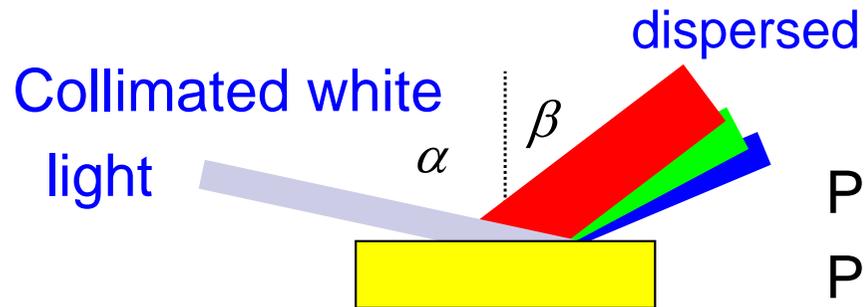
m : diffraction order

(conversion of wavelength to angle)

2.2. energy resolution - dispersion and focus -

How can we **monochromatize** by using a **diffraction grating**?

Most **basic** mode: **collimated-light illumination**



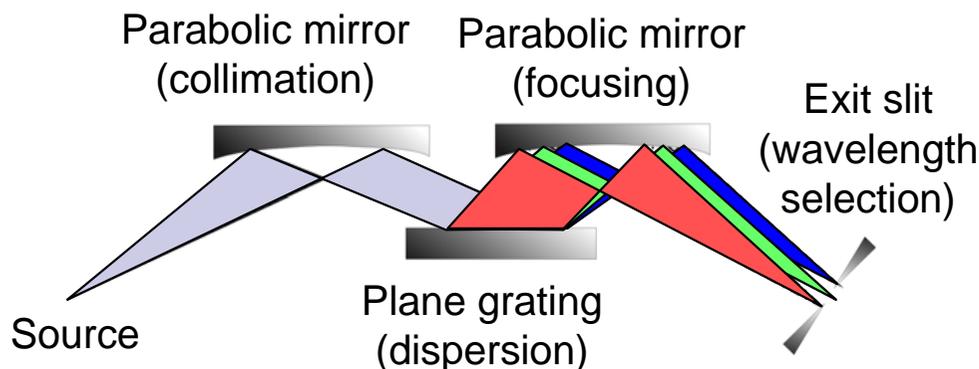
$$\sin \alpha + \sin \beta = nm\lambda$$

Problem 1: SR is **not a collimated light** !

Problem 2: **Superposition** of diffracted lights
 \Rightarrow **difficult to be resolved**

Solution 1: **Collimation** of diverging light with a **parabolic mirror**

Solution 2: **Focusing** of diffracted lights with another **parabolic mirror**

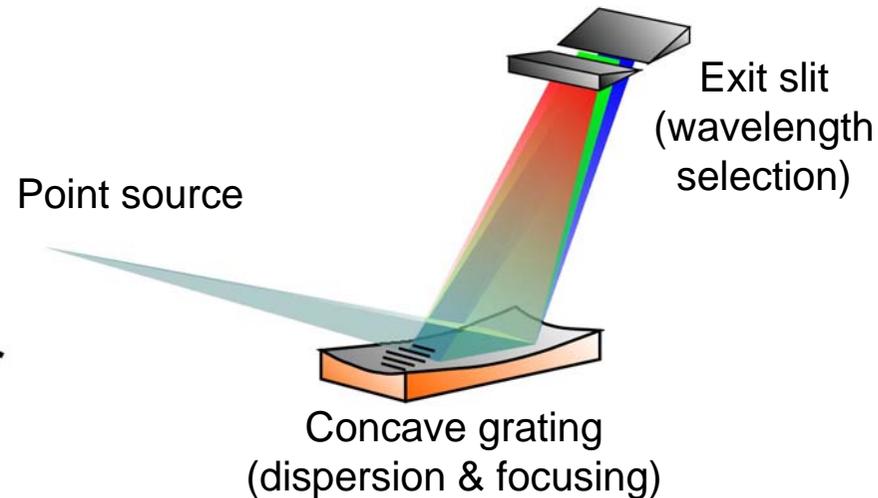


Focused diffracted lights are **well resolved** in wavelength at the exit slit !

Dispersion and Focus

2.2. energy resolution - dispersion and focus -

The **simplest** monochromator



Both the “**dispersion**” and “**focus**” are achieved by a **diffraction grating** only.

Is that really possible?

It is impossible to obtain a perfect focus at all wavelength

But, “perfect focus” is not necessary !

Small number of optical elements

2.2. energy resolution - dispersion and focus -

What determines the energy resolution?

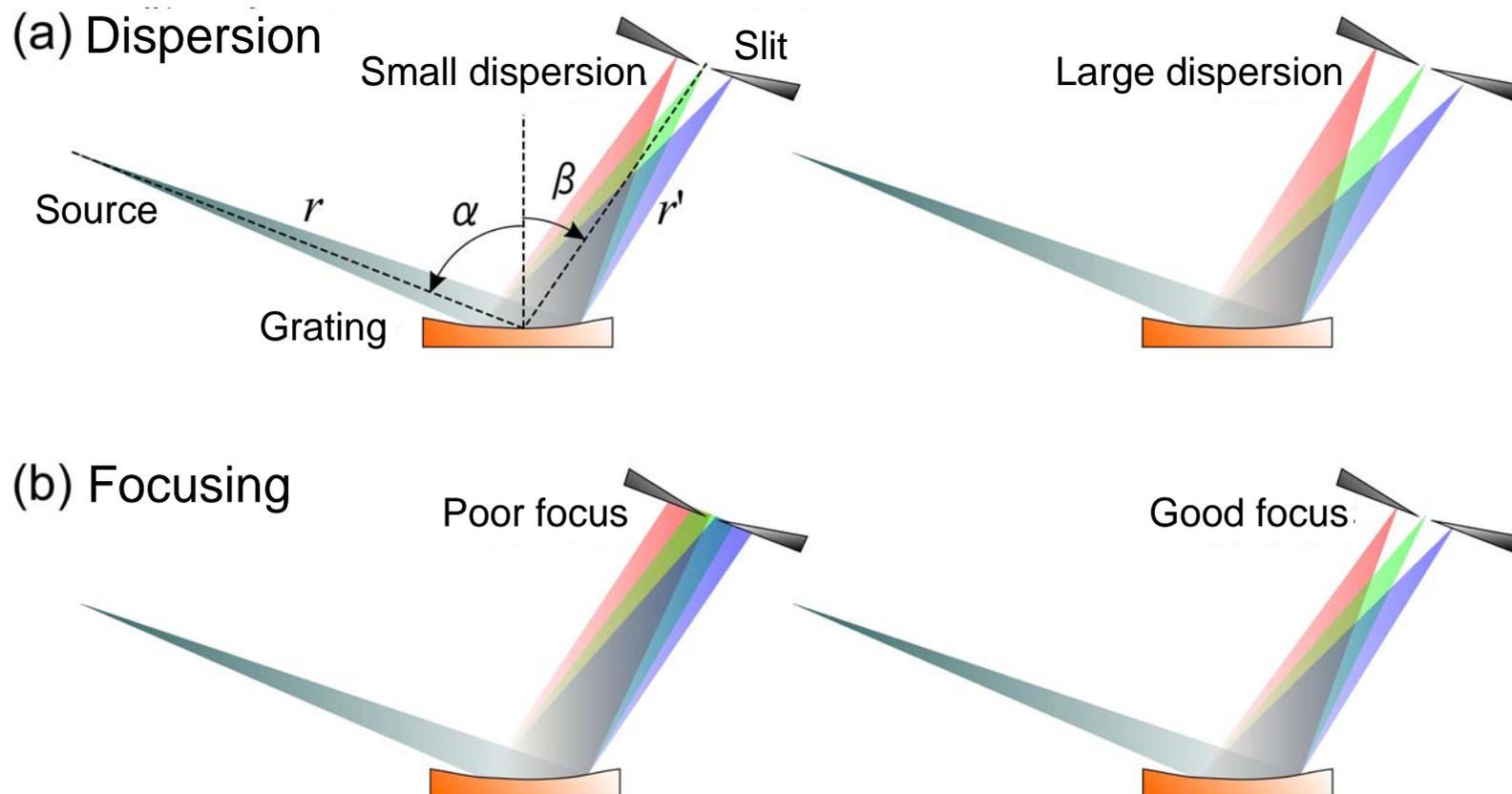
“Dispersion”: separation of lights with different wavelengths $\Rightarrow dz/d\lambda$

$$\sin \alpha + \sin \beta = nm\lambda \Rightarrow$$

$$\cos \beta d\beta = nmd\lambda, \quad dz/d\lambda = r' d\beta / d\lambda = r' nm / \cos \beta$$

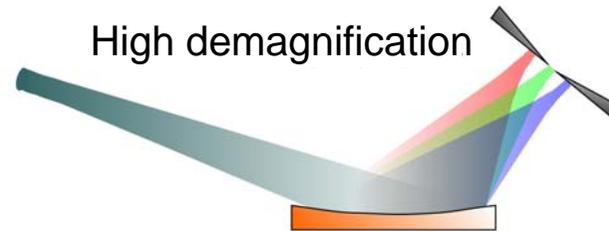
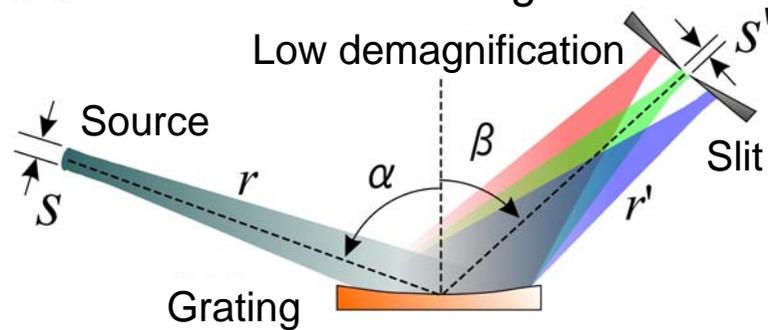
Ratio between “dispersion” and “light size” determines resolution.

i.e. large dispersion & small focus \Rightarrow high energy resolution

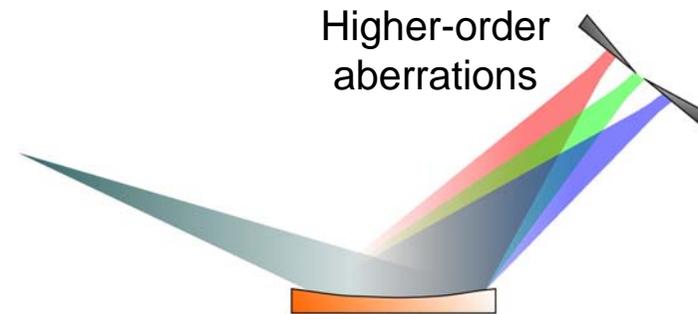
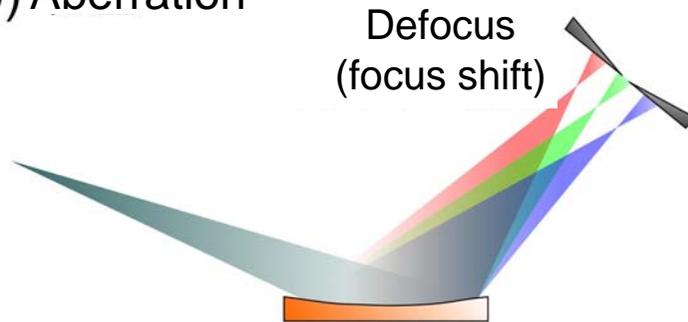


2.2. energy resolution - dispersion and focus - What determines the focus size?

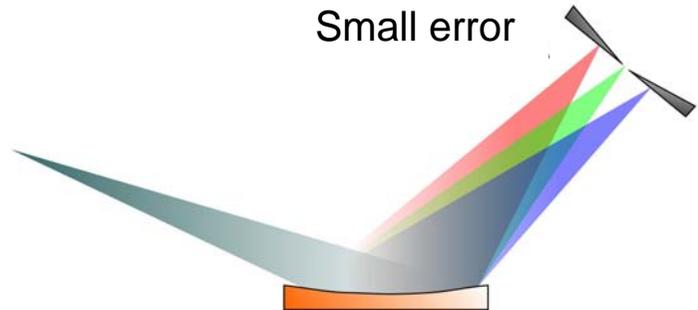
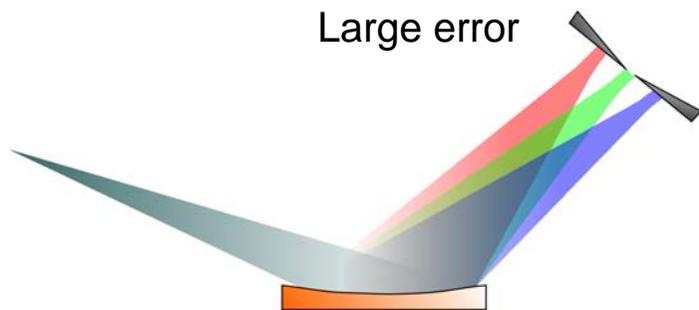
(a) Source size & Demagnification



(b) Aberration



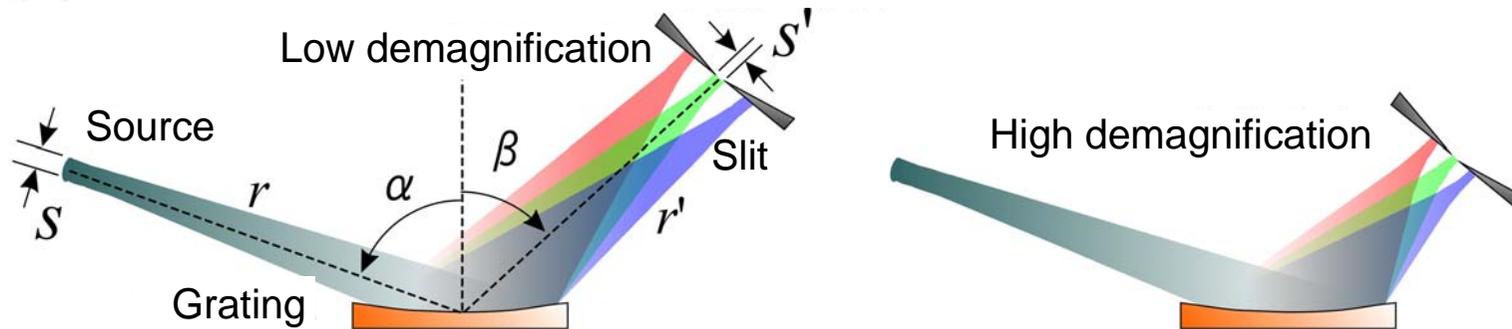
(c) Slope error (figure error)



2.2. energy resolution - dispersion and focus -

(a) Source size & Demagnification

(a) Source size & Demagnification



$$s' \text{ (lower limit)} = s d/d'$$

d : **divergence** at the source

d' : **divergence** at the focus

How to **reduce** s' :

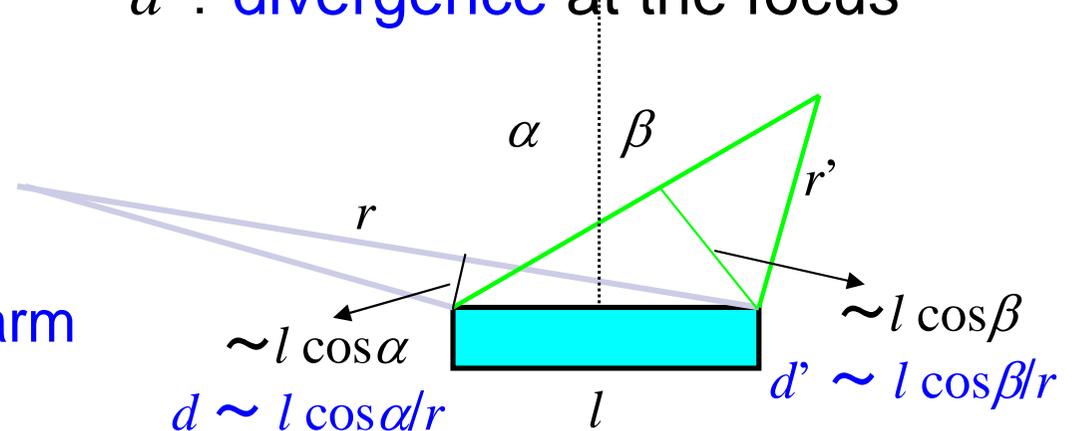
1. reduce r' compared to r

long entrance arm & short exit arm

2. decrease $|\beta|$ compared to α

make the incidence angle more grazing

(keeping the diffraction condition satisfied)



\Rightarrow **High groove density**

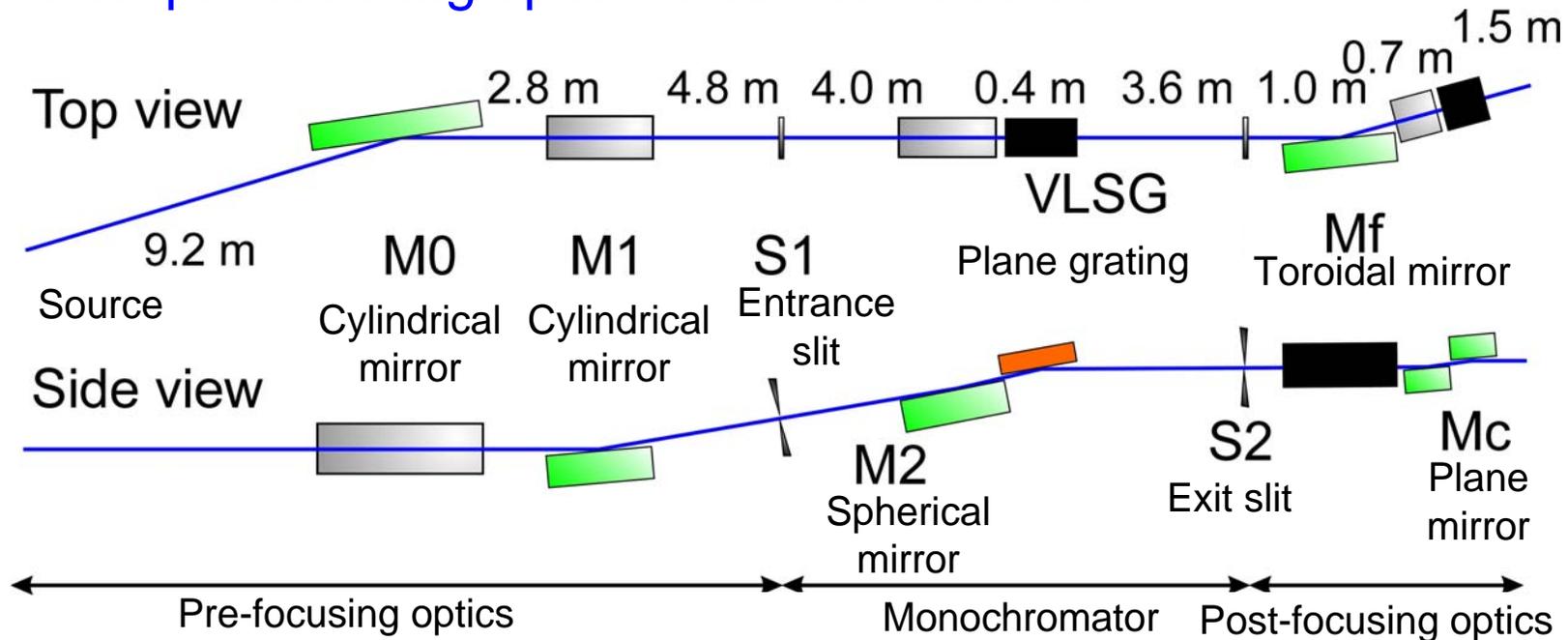
large included angle $(\alpha - \beta)$

2.2. energy resolution - dispersion and focus -

3. reduce the source size (s)

Must we reduce the size of the SR source itself?

Use of “pre-focusing optics” and “entrance slit”



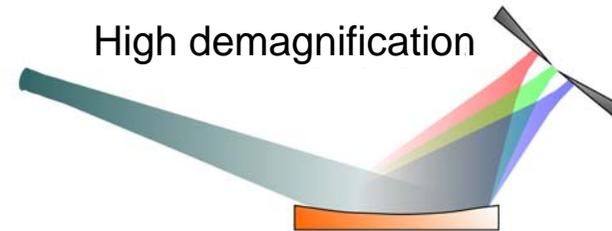
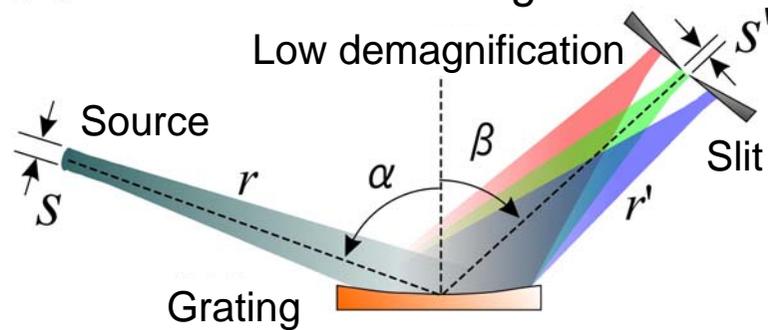
“Entrance slit” can be regarded as a **virtual source**

Source size can be controlled by entrance-slit opening
(at the sacrifice of intensity)

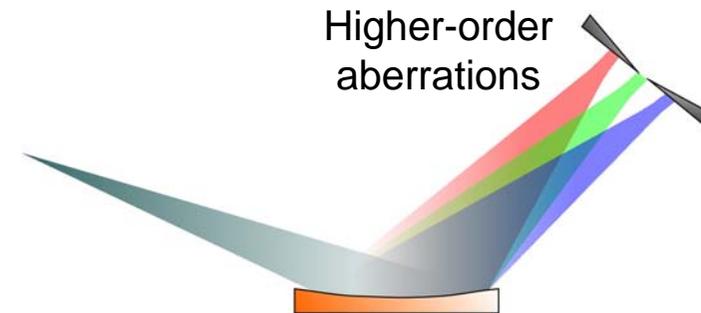
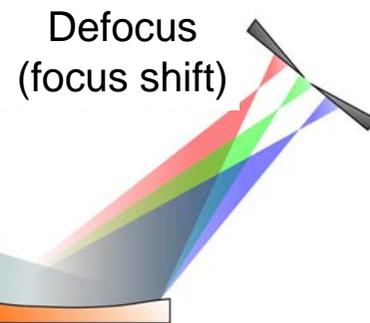
Demagnification in the pre-focusing optics is effective
(* divergence increases)

2.2. energy resolution - dispersion and focus - What determines the focus size?

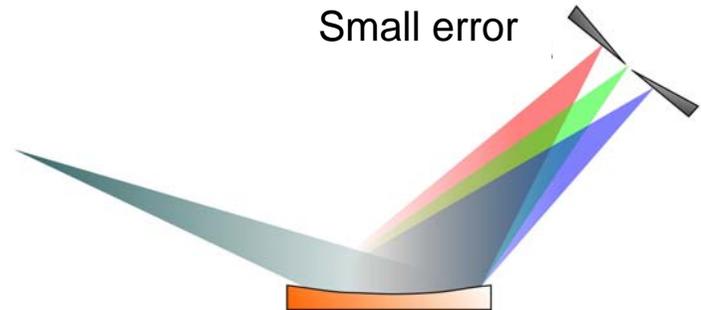
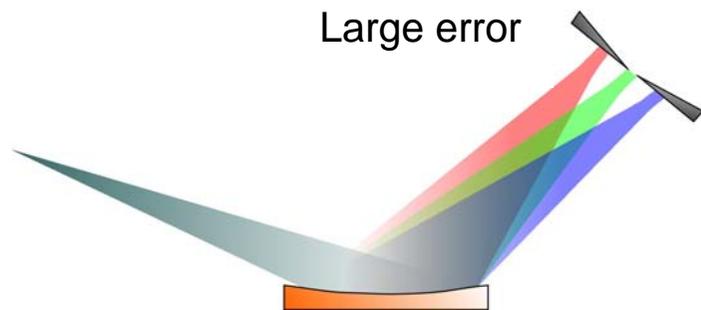
(a) Source size & Demagnification



(b) Aberration



(c) Slope error (figure error)



2.2. energy resolution - dispersion and focus -

(b) Aberration

Caused by a deviation from the elliptical (or parabolic) shape

e.g. Use of a spherical mirror instead of elliptical one

Diffraction effects should be taken into account for a diffraction grating

Aberration is usually expanded in a power series of the

position on the optical element, (w , l), using the **light path function**, F

$$F = p_A + q_A + r_B$$

$$+ M_{10}w + (M_{20}w^2 + M_{02}l^2 + M_{30}w^3 + M_{12}wl^2)/2$$

$$+ (M_{40}w^4 + M_{22}w^2l^2 + M_{04}l^4)/8 + \dots$$

$$+ [n_{10}w + (n_{20}w^2)/2 + (n_{30}w^3)/2 + (n_{40}w^4)/8 + \dots] m\lambda, \quad \leftarrow \text{For grating}$$

Diffraction condition

Defocus

(deviation from focal condition)

“coma” aberration

$$M_{10} = -\sin \alpha - \sin \beta,$$

$$M_{20} = (\cos^2 \alpha)/r_A + (\cos^2 \beta)/r_B,$$

$$M_{30} = (\sin \alpha \cos^2 \alpha)/r_A^2 + (\sin \beta \cos^2 \beta)/r_B^2$$

(sagittal focusing) or

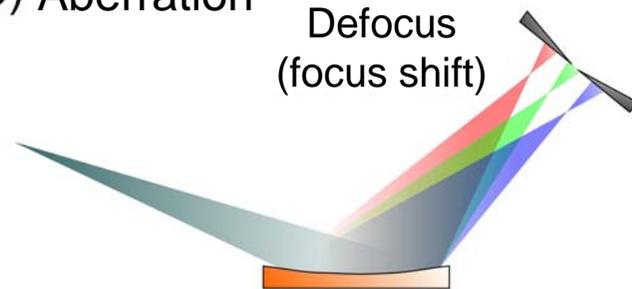
$$M_{30} = (\sin \alpha \cos^2 \alpha)/r_A^2 + (\sin \beta \cos^2 \beta)/r_B^2 - [2(A_{10})_A^2 K_A]/R_A$$

(meridional focusing),

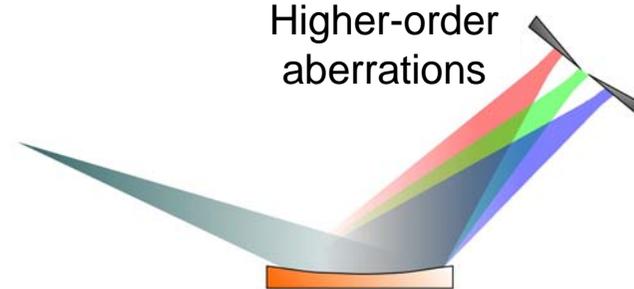
Larger illumination area (larger w and l) \Rightarrow larger effects of aberration 23

2.2. energy resolution - dispersion and focus -

(b) Aberration



Higher-order aberrations



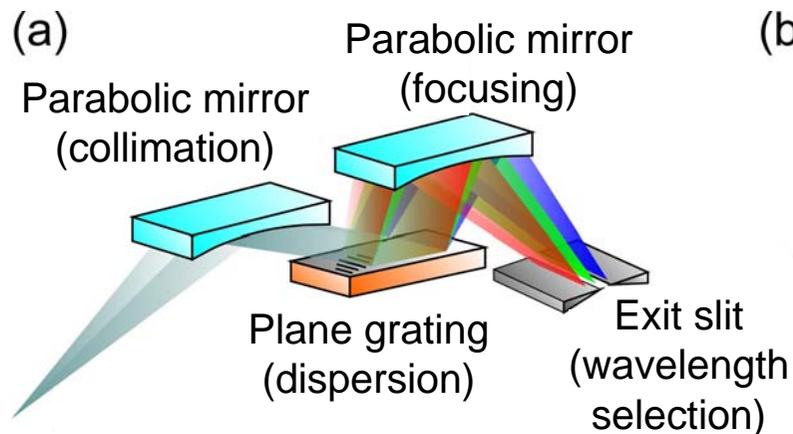
How to reduce the aberration:

Defocus can be compensated by adjusting the exit-slit position

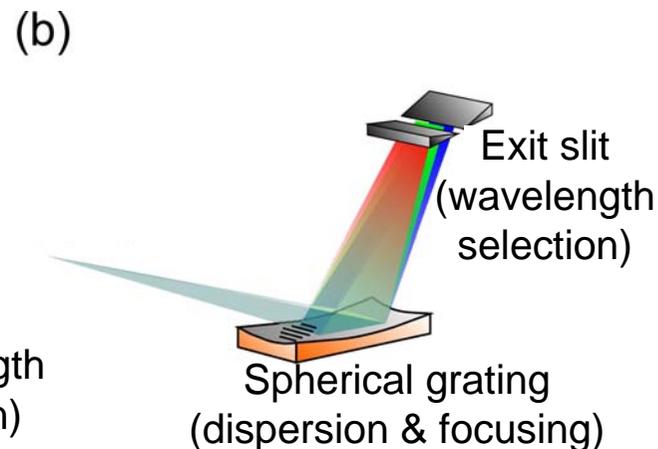
Higher-order aberrations can be canceled by a combination of mirrors (not easy)

Reduce the illumination area \Rightarrow small divergence (acceptance)

Examples for aberration-free or small-aberration optics



No aberration

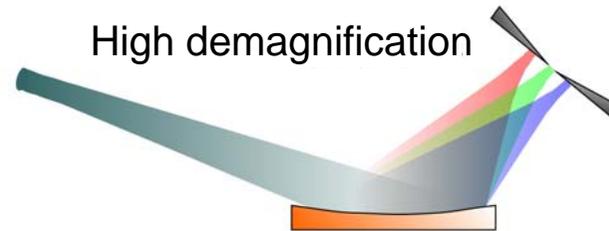
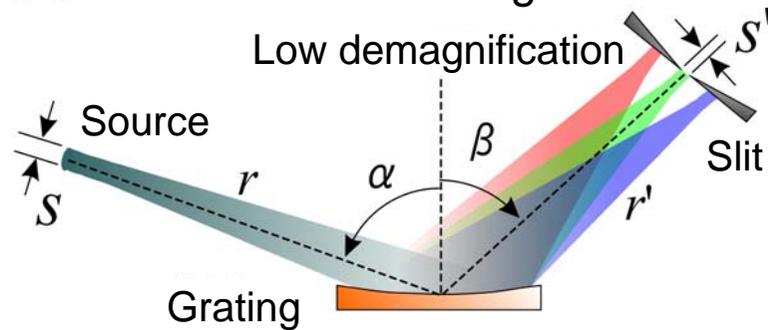


“Rowland condition”

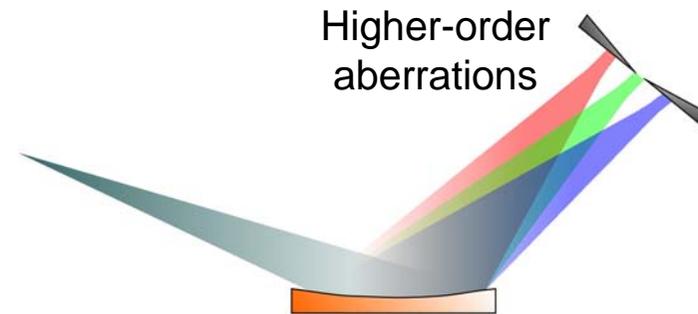
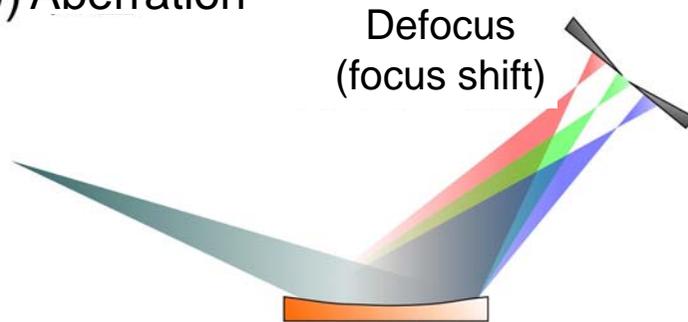
$$r = R \cos \alpha, r' = R \sin \beta$$

2.2. energy resolution - dispersion and focus - What determines the focus size?

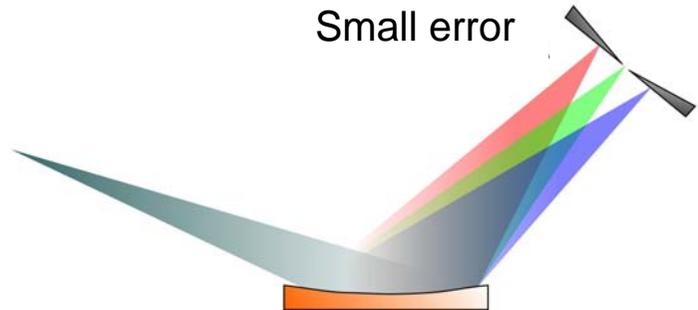
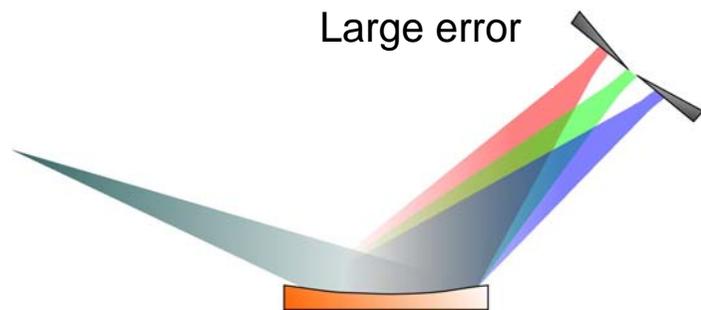
(a) Source size & Demagnification



(b) Aberration



(c) Slope error (figure error)



2.2. energy resolution - dispersion and focus -

(c) Slope errors

Errors in fabrication of optical elements (e.g. undulation of a plane mirror)

Not systematic \Rightarrow compensation is impossible

\Rightarrow One have to fabricate optical elements with small slope error

or design optics to reduce the effects of the slope errors

Demagnification is also effective to reduce the slope-error effects.

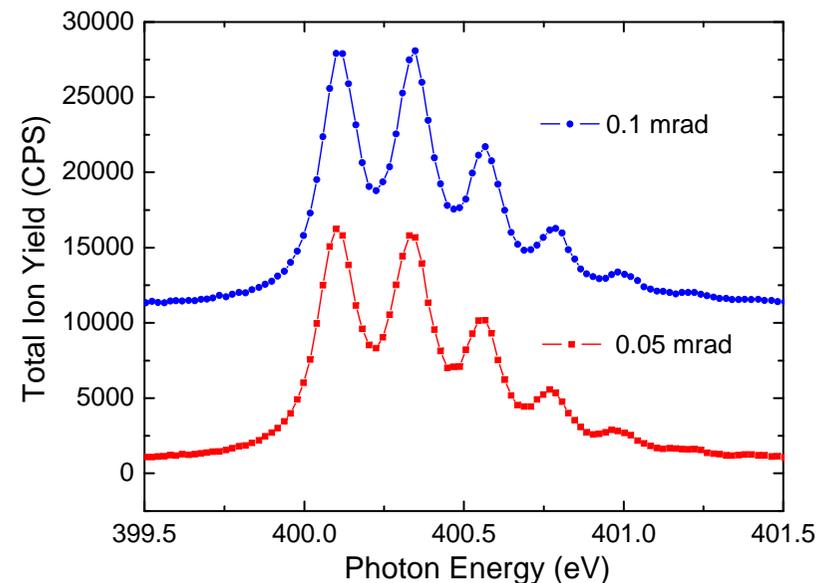
(d) Number of illuminated grooves

Intrinsic problem of diffraction

Resolving power $(\lambda/\Delta\lambda) \sim N$

* Small divergence

\Rightarrow small effects of aberration but small number of grooves



2.3. fabrication of gratings

1. Substrate (Plane, Cylindrical, Spherical, Toroidal,...)

Same as mirror fabrication

2. Fabrication of grooves (uniform or varied line spacing)

$$N = N_0 (1 + a_1 w + a_2 w^2 + a_3 w^3) \quad (\text{groove density})$$

$$F = p_A + q_A + r_B$$

$$+ M_{10}w + (M_{20}w^2 + M_{02}l^2 + M_{30}w^3 + M_{12}wl^2)/2$$

$$+ (M_{40}w^4 + M_{22}w^2l^2 + M_{04}l^4)/8 + \dots$$

$$+ [n_{10}w + (n_{20}w^2)/2 + (n_{30}w^3)/2 + (n_{40}w^4)/8 + \dots] m\lambda,$$

$$= N_0$$

= 0 for uniform line spacing

(a) Mechanical ruling

All groove parameters (a_1, a_2, \dots) can be controlled !

Relatively rough surface \Rightarrow causes stray light

Suitable for “Brazed” gratings

2.3. fabrication of gratings

(b) Holographic recording

Interference patterns of Laser lights

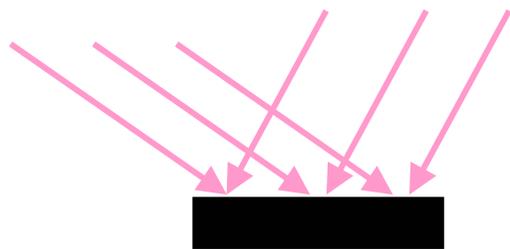
Some groove parameters might not be controlled

* aspheric wavefront recording is available

Relatively smooth surface \Rightarrow high reflectivity & low stray light

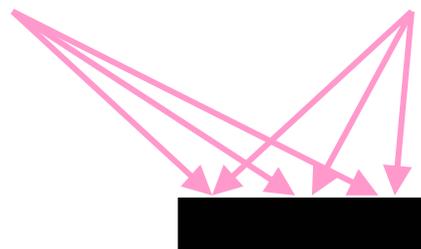
Both the “Laminar” and “Brazeed” gratings can be fabricated

* some manufacturer strongly prefers the Laminar type



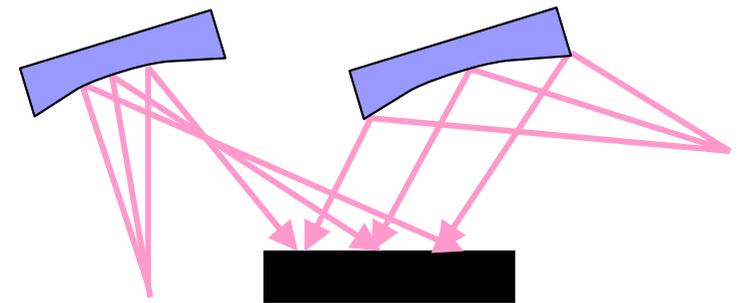
Collimated lights

\Rightarrow uniform line spacing



Spherical wavefronts

\Rightarrow varied line spacing
(poor control)



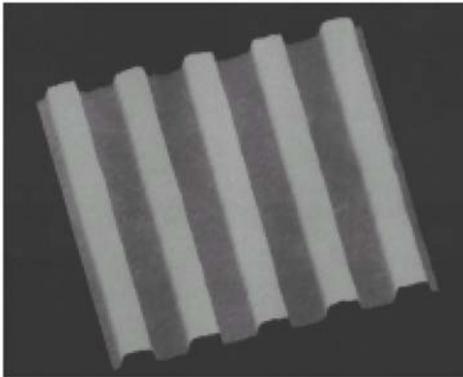
Aspheric wavefronts

\Rightarrow varied line spacing
(fine control)

2.3. fabrication of gratings

3. Groove shape (Laminar & Brazed)

(a) Laminar type

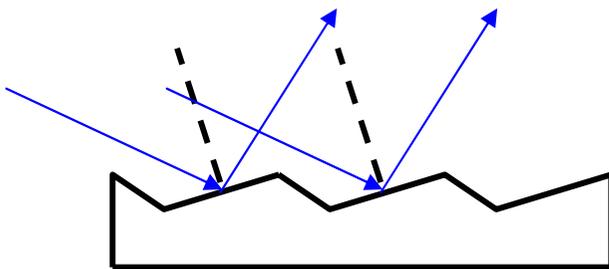


Medium diffraction efficiency (typically 20-30%)

Higher order suppression

interference between top and bottom parts

(b) Brazed type



High diffraction efficiency when “on Braze”

Strong higher orders



Thank you for your attention !